## Lecture 19

# Radiative transitions

#### Radiative transitions: background

- Previously, we have formulated a quantum theory of atoms (matter) coupling to a classical time-independent electromagnetic field, cf. Zeeman and Stark effects.
- To develop a fully quantum theory of light-matter systems, we have to address both the quantum theory of the electromagnetic field and formulate a theory of the coupling of light to matter.
- In the following we will address each of these components in turn, starting with light-matter coupling.
- Our motivation for developing such a consistent theory is that it:
  - (a) provides a platform to study radiative tranistions in atoms (which will address)
  - (b) forms the basis of **quantum optics** (which will not address but which is well-represented in subsequent courses).

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- Coupling of matter to electromagnetic field
- Spontaneous emission, absorption and stimulated emission
- Einstein's A and B coefficients
- Selection rules
- Theory of the laser and coherent states

• For a single-electron atom in a time-dependent external EM field, the Hamiltonian takes the form,

$$\hat{H}_{\mathrm{atom}} = rac{1}{2m} \left( \hat{\mathbf{p}} + e\mathbf{A}(\mathbf{r}, t) \right)^2 - e\phi(\mathbf{r}, t) + V(\mathbf{r})$$

(with a straightforward generalization to multi-electron atoms).

• Previously, we have seen that it is profitable to expand Hamiltonian in  $\hat{A}$ ,  $\hat{H}_{atom} = \hat{H}_0 + \hat{H}_{para} + \hat{H}_{dia.}$ , where  $\hat{H}_0 = \frac{\hat{\mathbf{p}}^2}{2m} + V(\mathbf{r}) - e\phi(t)$ 

$$\hat{H}_{ ext{para}}(t) = rac{e}{m} \mathbf{A}(t) \cdot \hat{\mathbf{p}}$$

the paramagnetic term describes coupling of the atom to the EM field, and  $\hat{H}_{\rm dia} = (e\mathbf{A})^2/2m$  represents diamagnetic term.

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• When quantized, EM field is described by the photon Hamiltonian,

$$\hat{H}_{
m rad} = \sum_{\mathbf{k},\lambda=1,2} \hbar \omega_{\mathbf{k}} \left( a_{\mathbf{k}\lambda}^{\dagger} a_{\mathbf{k}\lambda} + rac{1}{2} 
ight), \qquad \omega_{\mathbf{k}} = c |\mathbf{k}|$$

where  $a_{\mathbf{k}\lambda}^{\dagger}/a_{\mathbf{k}\lambda}$  create/annihilate photons with polarization  $\lambda$ .

These operators obey (bosonic) commutation relations,

$$[a_{\mathbf{k}\lambda}, a^{\dagger}_{\mathbf{k}'\lambda'}] = \delta_{\mathbf{k},\mathbf{k}'}\delta_{\lambda,\lambda'}, \qquad [a_{\mathbf{k}\lambda}, a_{\mathbf{k}'\lambda'}] = [a^{\dagger}_{\mathbf{k}\lambda}, a^{\dagger}_{\mathbf{k}'\lambda'}] = 0$$

and act on photon number states,  $|n_{k\lambda}\rangle = \frac{1}{\sqrt{n_{k\lambda}!}} (a_{k\lambda}^{\dagger})^{n_{k\lambda}} |\Omega\rangle$ , as

$$a_{\mathbf{k}\lambda}|n_{\mathbf{k}\lambda}\rangle = \sqrt{n_{\mathbf{k}\lambda}}|n_{\mathbf{k}\lambda}-1\rangle, \qquad a_{\mathbf{k}\lambda}^{\dagger}|n_{\mathbf{k}\lambda}\rangle = \sqrt{n_{\mathbf{k}\lambda}+1}|n_{\mathbf{k}\lambda}+1\rangle$$

With these definitions, the vector potential is given by

$$\hat{\mathbf{A}}(\mathbf{r}) = \sum_{\mathbf{k}\lambda=1,2} \sqrt{\frac{\hbar}{2\epsilon_0 \omega_{\mathbf{k}} V}} \left[ \hat{\mathbf{e}}_{\mathbf{k}\lambda} a_{\mathbf{k}\lambda} e^{i\mathbf{k}\cdot\mathbf{r}} + \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* a_{\mathbf{k}\lambda}^{\dagger} e^{-i\mathbf{k}\cdot\mathbf{r}} \right]$$

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## Heisenberg representation

$$\hat{H}_{\rm rad} = \sum_{\mathbf{k}\lambda} \hbar \omega_{\mathbf{k}} \left( a_{\mathbf{k}\lambda}^{\dagger} a_{\mathbf{k}\lambda} + \frac{1}{2} \right)$$

- To determine radiative transition rates, we will exploit Fermi's Golden rule. To prepare for this, it is convenient to transfer time-dependence to operators (Heisenberg representation).
- As with any operator, the field operators obey equations of motion,

$$\dot{a}_{k\lambda} = \frac{i}{\hbar} [\hat{H}, a_{k\lambda}] = i\omega_{k} [a_{k\lambda}^{\dagger} a_{k\lambda}, a_{k\lambda}] = -i\omega_{k} a_{k\lambda}$$

i.e.  $a_{k\lambda}(t) = a_{k\lambda}(0)e^{-i\omega_k t}$ ,  $a^{\dagger}_{k\lambda}(t) = a^{\dagger}_{k\lambda}(0)e^{i\omega_k t}$  and

$$\hat{\mathbf{A}}(\mathbf{r},t) = \sum_{\mathbf{k}\lambda} \sqrt{\frac{\hbar}{2\epsilon_0 \omega_{\mathbf{k}} V}} \left[ \hat{\mathbf{e}}_{\mathbf{k}\lambda} a_{\mathbf{k}\lambda} e^{i(\mathbf{k}\cdot\mathbf{r}-\omega_{\mathbf{k}}t)} + \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* a_{\mathbf{k}\lambda}^{\dagger} e^{-i(\mathbf{k}\cdot\mathbf{r}-\omega_{\mathbf{k}}t)} \right]$$

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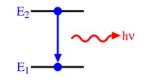
• Putting together all of these components, the total Hamiltonian is then given by  $\hat{H} = \hat{H}_{atom} + \hat{H}_{para} + \hat{H}_{rad}$  where (with  $\phi = 0$ )

$$\begin{split} \hat{H}_{\rm atom} &= \frac{\hat{\mathbf{p}}^2}{2m} + V(\mathbf{r}), \quad \hat{H}_{\rm rad} = \sum_{\mathbf{k}\lambda} \hbar \omega_{\mathbf{k}} \, \left( a_{\mathbf{k}\lambda}^{\dagger} a_{\mathbf{k}\lambda} + \frac{1}{2} \right) \quad \text{and} \\ \hat{H}_{\rm para}(t) &= \frac{e}{m} \hat{\mathbf{A}}(\mathbf{r}, t) \cdot \hat{\mathbf{p}} \quad \text{with} \end{split}$$

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 In the following, we will apply this Hamiltonian to the problem of radiative transitions in single electron atoms.

• Consider probability for an atom, initially in state  $|i\rangle$  to make transition to  $|f\rangle$  with emission of a photon of wavevector **k** and polarization  $\lambda$  – spontaneous emission.

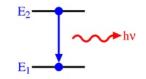


- If radiation field initially prepared in vacuum state, |Ω⟩, then final state involves one photon, a<sup>†</sup><sub>kλ</sub>|Ω⟩.
- Therefore, making use of Fermi's Golden rule, with the perturbation

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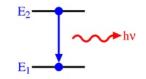


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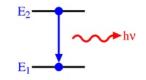


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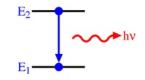


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$$\Gamma_{\mathbf{i}\to\mathbf{f},\mathbf{k}\lambda} = \frac{2\pi}{\hbar} \left| \langle \mathbf{f} | \otimes \langle \Omega | \mathbf{a}_{\mathbf{k}\lambda} \frac{e}{m} \sqrt{\frac{\hbar}{2\epsilon_0 \omega_{\mathbf{k}} V}} \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \mathbf{a}_{\mathbf{k}\lambda}^{\dagger} e^{-i\mathbf{k}\cdot\mathbf{r}} \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle \otimes |\Omega\rangle \right|^2 \delta(\mathbf{E}_{\mathbf{i}} - \mathbf{E}_{\mathbf{f}} - \hbar\omega_{\mathbf{k}})$$

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- To determine transition rate, we must analyse matrix elements of the form ⟨f|e<sup>-ik·r</sup>ê<sup>\*</sup><sub>kλ</sub> · β̂|i⟩. For typical state, ⟨ê<sup>\*</sup><sub>kλ</sub> · β̂⟩ ~ p ~ Zmcα.
- But what about exponential factor? With  $r \sim \hbar/p \simeq \hbar/mZc\alpha$ , and  $\omega_{\mathbf{k}} = c|\mathbf{k}| \sim \frac{\rho^2}{2m}$  (for electronic transitions), we have

$$\mathbf{k} \cdot \mathbf{r} \simeq rac{\omega_{\mathbf{k}}}{c} rac{\hbar}{p} \simeq rac{\hbar p}{mc} \simeq Z lpha$$

i.e. for  $Z\alpha \ll 1$ , we can expand exponential as power series in  ${\bf k} \cdot {\bf r}$  with lowest terms dominant.

• Taking zeroth order term, and using  $\hat{\mathbf{p}} = \frac{im}{\hbar} [\hat{H}_0, \mathbf{r}]$  (cf. Ehrenfest)

$$\langle \mathbf{f} | \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle = m \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \cdot \langle \mathbf{f} | \frac{i}{\hbar} [\hat{H}_0, \mathbf{r}] | \mathbf{i} \rangle = im \frac{E_{\mathrm{f}} - E_{\mathrm{i}}}{\hbar} \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \cdot \langle \mathbf{f} | \mathbf{r} | \mathbf{i} \rangle$$

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$$\langle \mathbf{f} | \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle = m \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \cdot \langle \mathbf{f} | \frac{i}{\hbar} [\hat{\mathcal{H}}_0, \mathbf{r}] | \mathbf{i} \rangle = \textit{im} \frac{\mathcal{E}_{\mathrm{f}} - \mathcal{E}_{\mathrm{i}}}{\hbar} \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \cdot \langle \mathbf{f} | \mathbf{r} | \mathbf{i} \rangle$$

$$\Gamma_{\mathbf{i}\to\mathbf{f},\mathbf{k}\lambda} = \frac{2\pi}{\hbar} \left| \langle \mathbf{f} | \frac{e}{m} \sqrt{\frac{\hbar}{2\epsilon_0 \omega_{\mathbf{k}} V}} e^{-i\mathbf{k}\cdot\mathbf{r}} \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle \right|^2 \delta(E_{\mathbf{i}} - E_{\mathbf{f}} - \hbar\omega_{\mathbf{k}})$$

- To determine transition rate, we must analyse matrix elements of the form ⟨f|e<sup>-ik·r</sup>ê<sup>\*</sup><sub>kλ</sub> · β̂|i⟩. For typical state, ⟨ê<sup>\*</sup><sub>kλ</sub> · β̂⟩ ~ p ~ Zmcα.
- But what about exponential factor? With  $r \sim \hbar/p \simeq \hbar/mZc\alpha$ , and  $\omega_{\mathbf{k}} = c|\mathbf{k}| \sim \frac{p^2}{2m}$  (for electronic transitions), we have

$$\mathbf{k} \cdot \mathbf{r} \simeq \frac{\omega_{\mathbf{k}}}{c} \frac{\hbar}{p} \simeq \frac{\hbar p}{mc} \simeq Z \alpha$$

i.e. for  $Z\alpha \ll 1$ , we can expand exponential as power series in  ${\bf k}\cdot {\bf r}$  with lowest terms dominant.

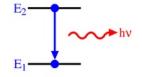
• Taking zeroth order term, and using  $\hat{\mathbf{p}} = \frac{im}{\hbar} [\hat{H}_0, \mathbf{r}]$  (cf. Ehrenfest)

 $\langle f | \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \cdot \hat{\mathbf{p}} | i \rangle = -im\omega_{\mathbf{k}} \langle f | \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* \cdot \mathbf{r} | i \rangle$ 

## Spontaneous emission: electric dipole approximation

$${}_{\lambda}\cdot\hat{\mathbf{p}}|\mathrm{i}
angle=-im\omega_{\mathbf{k}}\langle\mathrm{f}|\hat{\mathbf{e}}_{\mathbf{k}\lambda}^{*}\cdot\mathbf{r}|\mathrm{i}
angle$$

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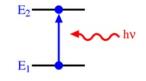


 This result, which emerges from leading approximation in Zα, is known as electric dipole approximation: Effectively, we have set

$$\hat{H}_{\text{para}} = \frac{e}{m}\hat{A}(\mathbf{r},t)\cdot\hat{\mathbf{p}} \simeq e\hat{\mathbf{E}}(\mathbf{r},t)\cdot\mathbf{r} = -\hat{\mathbf{E}}(\mathbf{r},t)\cdot\hat{\mathbf{d}}$$

translating to the potential energy of a dipole, with moment  $\hat{\mathbf{d}} = -e\mathbf{r}$ , in an oscillating electric field.

• Consider now absorption of a photon. If we assume that, in the initial state, there are  $n_{k\lambda}$  photons in mode  $(k\lambda)$  then, after the transition, there will be  $n_{k\lambda} - 1$  photons.



• Then, if initial state of the atom is  $|i\rangle$  and final state is  $|f\rangle$ ,

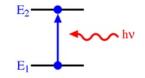
$$\begin{split} \mathbf{f} &| \otimes \langle (n_{\mathbf{k},\lambda} - 1) | \hat{H}_{\text{para}} | \mathbf{i} \rangle \otimes | n_{\mathbf{k}\lambda} \rangle \\ &= \langle \mathbf{f} | \otimes \langle (n_{\mathbf{k},\lambda} - 1) | \frac{e}{m} \sqrt{\frac{\hbar}{2\epsilon_0 \omega_{\mathbf{k}} V}} \hat{\mathbf{e}}_{\mathbf{k}\lambda} a_{\mathbf{k}\lambda} e^{i\mathbf{k}\cdot\mathbf{r}} \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle \otimes | n_{\mathbf{k}\lambda} \rangle \end{split}$$

• Then, using the relation  $a_{\mathbf{k}\lambda}|n_{\mathbf{k}\lambda}
angle = \sqrt{n_{\mathbf{k}\lambda}}|(n_{\mathbf{k}\lambda}-1)
angle$ ,

$$\langle \mathbf{f} | \otimes \langle (\mathbf{n}_{\mathbf{k},\lambda} - 1) | \hat{H}_{\text{para}} | \mathbf{i} \rangle \otimes | \mathbf{n}_{\mathbf{k}\lambda} \rangle = \langle \mathbf{f} | \frac{e}{m} \sqrt{\frac{\hbar n_{\mathbf{k}\lambda}}{2\epsilon_0 \omega_{\mathbf{k}} V}} e^{i\mathbf{k}\cdot\mathbf{r}} \hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle$$

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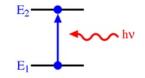
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• Then, using the relation  $a_{k\lambda}|n_{k\lambda}\rangle = \sqrt{n_{k\lambda}}|(n_{k\lambda}-1)\rangle$ ,

$$\langle \mathbf{f} | \otimes \langle (\mathbf{n}_{\mathbf{k},\lambda} - 1) | \hat{H}_{\text{para}} | \mathbf{i} \rangle \otimes | \mathbf{n}_{\mathbf{k}\lambda} \rangle = \langle \mathbf{f} | \frac{e}{m} \sqrt{\frac{\hbar n_{\mathbf{k}\lambda}}{2\epsilon_0 \omega_{\mathbf{k}} V}} e^{i\mathbf{k}\cdot\mathbf{r}} \hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle$$

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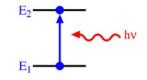
• Then, if initial state of the atom is  $|i\rangle$  and final state is  $|f\rangle$ ,

$$\begin{split} \mathrm{f} &| \otimes \langle (n_{\mathbf{k},\lambda} - 1) | \hat{\mathcal{H}}_{\mathrm{para}} | \mathrm{i} \rangle \otimes | n_{\mathbf{k}\lambda} \rangle \\ &= \langle \mathrm{f} | \otimes \langle (n_{\mathbf{k},\lambda} - 1) | \frac{e}{m} \sqrt{\frac{\hbar}{2\epsilon_0 \omega_{\mathbf{k}} V}} \hat{\mathbf{e}}_{\mathbf{k}\lambda} a_{\mathbf{k}\lambda} e^{i\mathbf{k}\cdot\mathbf{r}} \cdot \hat{\mathbf{p}} | \mathrm{i} \rangle \otimes | n_{\mathbf{k}\lambda} \rangle \end{split}$$

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$$\langle \mathbf{f} | \otimes \langle (\mathbf{n}_{\mathbf{k},\lambda} - \mathbf{1}) | \hat{H}_{\text{para}} | \mathbf{i} \rangle \otimes | \mathbf{n}_{\mathbf{k}\lambda} \rangle = \langle \mathbf{f} | \frac{e}{m} \sqrt{\frac{\hbar n_{\mathbf{k}\lambda}}{2\epsilon_0 \omega_{\mathbf{k}} V}} e^{i\mathbf{k}\cdot\mathbf{r}} \hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle$$

Consider now absorption of a photon. If we assume that, in the initial state, there are n<sub>kλ</sub> photons in mode (kλ) then, after the transition, there will be n<sub>kλ</sub> - 1 photons.



• As a result, using Fermi's Golden rule,

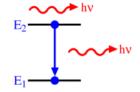
$$\Gamma_{\mathrm{i}
ightarrow \mathrm{f}}(t) = rac{2\pi}{\hbar^2} |\langle \mathrm{f}| \otimes \langle (\textit{\textit{n}}_{\mathbf{k}\lambda} - 1) | \hat{H}_{\mathrm{para}} | \mathrm{i} 
angle \otimes |\textit{\textit{n}}_{\mathbf{k}\lambda} 
angle |^2 \delta(\omega_{\mathrm{fi}} - \omega)$$

we obtain the transition amplitude,

$$\Gamma_{\mathbf{i}\to\mathbf{f},\mathbf{k}\lambda} = \frac{2\pi}{\hbar} \left| \langle \mathbf{f} | \frac{e}{m} \sqrt{\frac{\hbar n_{\mathbf{k}\lambda}}{2\epsilon_0 \omega_{\mathbf{k}} V}} e^{i\mathbf{k}\cdot\mathbf{r}} \hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle \right|^2 \delta(E_{\mathbf{f}} - E_{\mathbf{i}} - \hbar \omega_{\mathbf{k}})$$

• In particular, we find that the absorption rate increases linearly with photon number,  $n_{k\lambda}$ .

 Similarly, if we consider emission process in which there are already n<sub>kλ</sub> photons in initial state,

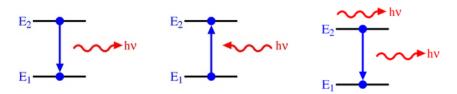


• using the relation  $a_{\mathbf{k}\lambda}^{\dagger}|n_{\mathbf{k}\lambda}\rangle = \sqrt{n_{\mathbf{k}\lambda}+1}|(n_{\mathbf{k}\lambda}+1)\rangle$ , we have revised transition rate,

$$\Gamma_{\mathbf{i}\to\mathbf{f},\mathbf{k}\lambda} = \frac{2\pi}{\hbar} \left| \langle \mathbf{f} | \frac{e}{m} \sqrt{\frac{\hbar(n_{\mathbf{k}\lambda}+1)}{2\epsilon_0 \omega_{\mathbf{k}} V}} e^{-i\mathbf{k}\cdot\mathbf{r}} \hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle \right|^2 \delta(E_{\mathbf{f}} - E_{\mathbf{i}} - \hbar\omega_{\mathbf{k}})$$

 Enhancement of transition rate by photon occupancy known as stimulated emission.

#### **Radiative transitions: summary**



• Altogether, in dipole approximation  $\langle f | \hat{\boldsymbol{e}}_{\boldsymbol{k}\lambda} \cdot \hat{\boldsymbol{p}} | i \rangle \simeq -im\omega_{\boldsymbol{k}} \langle f | \hat{\boldsymbol{e}}_{\boldsymbol{k}\lambda} \cdot \boldsymbol{r} | i \rangle$ 

$$\Gamma_{\mathbf{i}\to\mathbf{f},\mathbf{k}\lambda} \simeq \frac{\pi\omega_{\mathbf{k}}}{\epsilon_{0}V} |\langle \mathbf{f}|\hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{d}}|\mathbf{i}\rangle|^{2} \begin{cases} n_{\mathbf{k}\lambda} \,\,\delta(E_{\mathbf{f}} - E_{\mathbf{i}} - \hbar\omega_{\mathbf{k}}) & \text{absorption} \\ (n_{\mathbf{k}\lambda} + 1) \,\,\delta(E_{\mathbf{i}} - E_{\mathbf{f}} - \hbar\omega_{\mathbf{k}}) & \text{emission} \end{cases}$$

where  $\hat{\mathbf{d}} = -e\mathbf{r}$  is electric dipole operator.

- If there are no photons present initially,  $\Gamma_{i\to f, \bm{k}\lambda}$  reduces to result for spontaneous emission.
- The coincidence of  $n_{k\lambda}$ -independent coefficients for absorption and emission coincide is known as detailed balance.

#### Absorption and stimulated emission

$$\Gamma_{\mathbf{i}\to\mathbf{f},\mathbf{k}\lambda} \simeq \frac{\pi\omega_{\mathbf{k}}}{\epsilon_{0}\nu} |\langle \mathbf{f}|\hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{d}}|\mathbf{i}\rangle|^{2} \begin{cases} n_{\mathbf{k}\lambda} \,\,\delta(\mathcal{E}_{\mathbf{f}} - \mathcal{E}_{\mathbf{i}} - \hbar\omega_{\mathbf{k}}) & \text{absorption} \\ (n_{\mathbf{k}\lambda} + 1) \,\,\delta(\mathcal{E}_{\mathbf{i}} - \mathcal{E}_{\mathbf{f}} - \hbar\omega_{\mathbf{k}}) & \text{emission} \end{cases}$$

 Integrated transition rate associated with a small solid angle dΩ in the direction k given by

$$dR_{\mathrm{i}\to\mathrm{f},\lambda} = \sum_{\mathbf{k}\in d\Omega} \Gamma_{\mathrm{i}\to\mathrm{f},\mathbf{k}\lambda} = d\Omega \, V \int \frac{k^2 dk}{(2\pi)^3} \Gamma_{\mathrm{i}\to\mathrm{f},\mathbf{k}\lambda}$$

 If we assume that the photon number, n<sub>kλ</sub> is isotropic, independent of angle Ω, using the dispersion relation ω<sub>k</sub> = ck, we obtain

$$\frac{dR_{\mathrm{i}\to\mathrm{f},\lambda}}{d\Omega} = \frac{V}{c^3} \int \frac{\omega^2 d\omega}{(2\pi)^3} \frac{\pi\omega}{\epsilon_0 V} \langle \mathrm{f} | \hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{d}} | \mathrm{i} \rangle |^2 \begin{cases} n_\lambda(\omega) \ \delta(E_\mathrm{f} - E_\mathrm{i} - \hbar\omega) \\ (n_\lambda(\omega) + 1) \ \delta(E_\mathrm{i} - E_\mathrm{f} - \hbar\omega) \end{cases}$$
  
where  $\hbar\omega = |E_\mathrm{f} - E_\mathrm{i}|.$ 

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• From this expression, we can obtain the power loss as  $P_{\lambda} = \hbar \omega R_{\lambda}$ .

#### Absorption and stimulated emission

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where  $\hbar \omega = |E_{\rm f} - E_{\rm i}|$ .

• From this expression, we can obtain the power loss as  $P_{\lambda} = \hbar \omega R_{\lambda}$ .

In fact, frequency dependence of spontaneous emission rate can be inferred using ingenious argument due to Einstein who showed that stimulated and spontaneous transitions must be related.

- Consider ensemble of atoms exposed to black-body radiation at temperature T. Let us consider transitions between states  $|\psi_j\rangle$  and  $|\psi_k\rangle$ , with  $E_k E_j = \hbar\omega$ .
- If number of atoms in two states given by n<sub>j</sub> and n<sub>k</sub>, transition rates per atom given by:

absorption  $j \rightarrow k$   $B_{j \rightarrow k} u(\omega)$ stimulated emission  $k \rightarrow j$   $B_{k \rightarrow j} u(\omega)$ spontaneous emission  $k \rightarrow j$   $A_{k \rightarrow j}(\omega)$ 

where  $u(\omega)$  denotes energy density of radiation.

• A and B are known as **Einstein's A and B coefficients**, and, as we have seen, are properties of atomic states.

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#### **Einstein's A and B coefficients**

$$egin{aligned} & ext{absorption} & j o k & B_{j o k} u(\omega) \ & ext{stimulated emission} & k o j & B_{k o j} u(\omega) \ & ext{spontaneous emission} & k o j & A_{k o j}(\omega) \end{aligned}$$

• In thermodynamic equilibrium the rates must balance, so that

$$n_k \left[ A_{k \to j}(\omega) + B_{k \to j} u(\omega) \right] = n_j B_{j \to k} u(\omega)$$

 At the same time, relative populations of two states given by Boltzmann factor,

$$\frac{n_j}{n_k} = \frac{e^{-E_j/k_{\rm B}T}}{e^{-E_k/k_{\rm B}T}} = e^{\hbar\omega/k_{\rm B}T}$$

Thus we have:

$$A_{k\to j}(\omega) = \left[B_{j\to k}e^{\hbar\omega/k_{\rm B}T} - B_{k\to j}\right]u(\omega)$$

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• For black-body, energy density  $u(\omega)$  set by Planck formula,

$$u(\omega) = \frac{\hbar\omega^3}{\pi^2 c^3} \bar{n}(\omega) = \frac{\hbar\omega^3}{\pi^2 c^3} \frac{1}{e^{\hbar\omega/k_{\rm B}T} - 1}$$

 Since A<sub>k→j</sub> is intrinsic (independent of temperature), T must cancel on right hand side, i.e.

$$B_{k \to j} = B_{j \to k}$$
 and  $A_{k \to j}(\omega) = B_{k \to j} \frac{\hbar \omega^3}{\pi^2 c^3}$ 

• So, A and B coefficients are related, and if we can calculate B coefficient for stimulated emission from Fermi's Golden rule, we can infer A, and vice versa.

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- Formulae for rates  $\Gamma_{i \to f, \mathbf{k}\lambda}$  show that radiative transitions will not occur between states  $|i\rangle$  and  $|f\rangle$  unless at least one component of the dipole matrix element  $\langle f | \hat{\mathbf{d}} | i \rangle$  is non-zero.
- If matrix elements are zero for certain pairs, they are disallowed (at least in the electric dipole approximation) leading to selection rules.
- Since dipole operator  $\hat{\mathbf{d}} = -e\mathbf{r}$  changes sign under parity  $(\mathbf{r} \rightarrow -\mathbf{r})$ , matrix element  $\langle f | \hat{\mathbf{d}} | i \rangle$  will vanish if  $| f \rangle$  and  $| i \rangle$  have same parity.

The parity of the wavefunction must change in an electric dipole transition.

$$\Gamma_{\mathbf{i}\to\mathbf{f},\mathbf{k}\lambda} \simeq \frac{\pi\omega_{\mathbf{k}}}{\epsilon_{0}V} |\langle \mathbf{f}|\hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{d}}|\mathbf{i}\rangle|^{2} \begin{cases} n_{\mathbf{k}\lambda} \ \delta(E_{\mathbf{f}} - E_{\mathbf{i}} - \hbar\omega_{\mathbf{k}}) & \text{absorption} \\ (n_{\mathbf{k}\lambda} + 1) \ \delta(E_{\mathbf{i}} - E_{\mathbf{f}} - \hbar\omega_{\mathbf{k}}) & \text{emission} \end{cases}$$

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• Separating wavefunction into spatial and spin components,  $|f\rangle = |\phi_f\rangle \otimes |\chi_f\rangle$ , since dipole operator acts only on spatial part,

$$\langle \mathrm{f}|\hat{\mathbf{d}}|\mathrm{i}
angle = -\langle \chi_\mathrm{f}|\chi_\mathrm{i}
angle \int d^3 r\, \phi^*_\mathrm{f}(\mathbf{r})\, e\mathbf{r}\phi_\mathrm{i}(\mathbf{r})$$

i.e. spin term,  $\langle\chi_f|\chi_i\rangle$  , vanishes unless  $|\chi_i\rangle$  and  $|\chi_f\rangle$  are identical,

$$\Delta s = 0, \qquad \Delta m_s = 0$$

**2** The spin state is not altered in an electric dipole transition.

$$\Gamma_{\mathbf{i}\to\mathbf{f},\mathbf{k}\lambda} \simeq \frac{\pi\omega_{\mathbf{k}}}{\epsilon_{0}V} |\langle \mathbf{f}|\hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{d}}|\mathbf{i}\rangle|^{2} \begin{cases} n_{\mathbf{k}\lambda} \ \delta(E_{\mathbf{f}} - E_{\mathbf{i}} - \hbar\omega_{\mathbf{k}}) & \text{absorption} \\ (n_{\mathbf{k}\lambda} + 1) \ \delta(E_{\mathbf{i}} - E_{\mathbf{f}} - \hbar\omega_{\mathbf{k}}) & \text{emission} \end{cases}$$

- From the operator identity,  $[\hat{L}_i, r_j] = i\hbar\epsilon_{ijk}r_k$ , it follows that  $[\hat{L}_z, z] = 0, \qquad [\hat{L}_z, x \pm iy] = \pm(x \pm iy)\hbar$
- We therefore obtain,

$$\langle \ell', m' | [\hat{L}_z, z] | \ell, m \rangle = (m' - m) \hbar \langle \ell', m' | z | \ell, m \rangle = 0$$

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$$[\tilde{L}_z, z] = 0, \qquad [\tilde{L}_z, x \pm iy] = \pm (x \pm iy)h$$

We therefore obtain,

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$$\langle \ell', m' | [\hat{\mathbf{L}}^2, [\hat{\mathbf{L}}^2, \mathbf{r}]] | \ell, m \rangle = [\ell'(\ell'+1) - \ell(\ell+1)]^2 \langle \ell', m' | \mathbf{r} | \ell, m \rangle$$
  
= 2[\left'(\left'+1) + \ell(\ell+1)] \left\left', m' | \mbox{\mathbf{r}} | \ell, m \rangle

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- One may summarize the selection rules for  $\ell$  and  $m_{\ell}$  is by saying that the photon carries off (or brings in, in an absorption transition) one unit of angular momentum.
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- Similarly, for electric dipole transitions with  $m' = m \pm 1$ ,  $\langle \ell', m' | x \mp i y | \ell, m \rangle = 0 = \langle \ell', m' | z | \ell, m \rangle$ , and  $\langle f | \mathbf{d} | i \rangle \sim (1, \mp i, 0)$ .

(a) If the wavevector of photon lies along z, the emitted light is circularly polarized with a polarization which depends on helicity.

(b) If the wavevector lies in *xy* place, the emitted light is linearly polarized, while in general it is elliptically polarized.

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## Selection rules: LS coupling

- In the presence of spin-orbit coupling, stationary states labelled by quantum numbers  $J, m_J, \ell, s$  where  $\hat{\mathbf{J}} = \hat{\mathbf{L}} + \hat{\mathbf{S}}$ .
- The selection rules in this case can be inferred by looking for the conditions for non-zero matrix elements (J', m<sub>J</sub>, l', s' |r|J, m<sub>J</sub>, l, s).
- By expanding states  $|J, m_J, \ell, s\rangle$  in basis states  $|\ell, m_\ell\rangle \otimes |s, m_s\rangle$ , one may uncover the following set of selection rules:

• For dipole transitions to take place, we require that  $\Delta m_j = 0, \pm 1$  $\Delta j = 0, \pm 1 \text{ not } 0 \to 0$ 

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## Radiative transitions: recap

• When coupled to a quantized electromagnetic field, the total Hamiltonian for atomic system given by  $\hat{H} = \hat{H}_{atom} + \hat{H}_{para} + \hat{H}_{rad}$  where

$$\hat{H}_{\mathrm{atom}} = \frac{\hat{\mathbf{p}}^2}{2m} + V(\mathbf{r}), \qquad \hat{H}_{\mathrm{rad}} = \sum_{\mathbf{k}\lambda} \hbar \omega_{\mathbf{k}} \left( a_{\mathbf{k}\lambda}^{\dagger} a_{\mathbf{k}\lambda} + \frac{1}{2} \right)$$

denotes the Hamiltonian of the isolated atomic and radiation field, and

$$\hat{H}_{\text{para}}(t) = \frac{e}{m}\hat{\mathbf{A}}(\mathbf{r},t)\cdot\hat{\mathbf{p}}$$

denotes the coupling with

$$\hat{\mathbf{A}}(\mathbf{r},t) = \sum_{\mathbf{k}\lambda} \sqrt{\frac{\hbar}{2\epsilon_0 \omega_{\mathbf{k}} V}} \left[ \hat{\mathbf{e}}_{\mathbf{k}\lambda} a_{\mathbf{k}\lambda} e^{i(\mathbf{k}\cdot\mathbf{r}-\omega_{\mathbf{k}}t)} + \hat{\mathbf{e}}_{\mathbf{k}\lambda}^* a_{\mathbf{k}\lambda}^{\dagger} e^{-i(\mathbf{k}\cdot\mathbf{r}-\omega_{\mathbf{k}}t)} \right]$$

## Radiative transitions: recap

• The transition rate between an initial and final state of the atom and electromagnetic field can be estimated using Fermi's Golden rule

$${\sf \Gamma}_{
m i 
ightarrow 
m f} = rac{2\pi}{\hbar^2} |\langle {
m f} | \hat{H}_{
m para} | {
m i} 
angle |^2 \delta(\omega_{
m if} - \omega)$$

where  $\hbar \omega_{if} = E_i - E_f$ .

- Crucially, since the photon creation/annihilation operators obey the relations,  $a_{\mathbf{k}\lambda}^{\dagger}|n_{\mathbf{k}\lambda}\rangle = \sqrt{n_{\mathbf{k}\lambda}+1}|(n_{\mathbf{k}\lambda}+1)\rangle$  and  $a_{\mathbf{k}\lambda}|n_{\mathbf{k}\lambda}\rangle = \sqrt{n_{\mathbf{k}\lambda}}|(n_{\mathbf{k}\lambda}-1)\rangle$  the transition rate depends on the photon number,  $n_{\mathbf{k}\lambda}$ .
- When  $Z\alpha \ll 1$ , the effective range of the interaction of the atom with the field is small (i.e.  $kr \sim Z\alpha$ ) and we can effect the dipole approximation,

$$\langle \mathbf{f} | e^{-i\mathbf{k}\cdot\mathbf{r}} \hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{p}} | \mathbf{i} \rangle \simeq \frac{im\omega_{\mathbf{k}}}{e} \langle \mathbf{f} | \hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \mathbf{d} | \mathbf{i} \rangle, \qquad \mathbf{d} = -e\mathbf{r}$$

## **Radiative transitions: recap**



• In the electric dipole approximation, the transition rate is given by

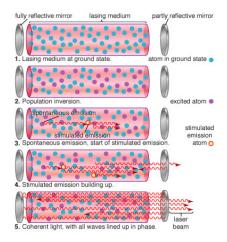
$$\Gamma_{\mathbf{i}\to\mathbf{f},\mathbf{k}\lambda} \simeq \frac{\pi\omega_{\mathbf{k}}}{\epsilon_{0}V} |\langle \mathbf{f}|\hat{\mathbf{e}}_{\mathbf{k}\lambda} \cdot \hat{\mathbf{d}}|\mathbf{i}\rangle|^{2} \begin{cases} n_{\mathbf{k}\lambda} \ \delta(E_{\mathbf{f}} - E_{\mathbf{i}} - \hbar\omega_{\mathbf{k}}) & \text{absorption} \\ (n_{\mathbf{k}\lambda} + 1) \ \delta(E_{\mathbf{i}} - E_{\mathbf{f}} - \hbar\omega_{\mathbf{k}}) & \text{emission} \end{cases}$$

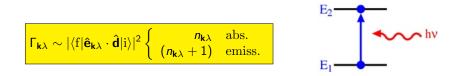
where  $\hat{\mathbf{d}} = -e\mathbf{r}$  is electric dipole operator.

- The coincidence of  $n_{k\lambda}$ -independent coefficients for absorption and emission coincide is known as detailed balance.
- From these results, we turn now to consider the principle of the operation of an atomic laser.

# Theory of laser

- Principle of stimulated emission provides basis of laser operation: light amplification by stimulated emission of radiation.
- However, laser not only amplifies light, but provides source of monochromatic (single mode), coherent (spatial/temporal), directional and intense radiation.
- In atomic laser, the gain medium provided by a gas of atoms confined to a cavity and bound by highly reflective mirrors.





- Consider gas of atoms in a cavity subject to an EM field of intensity  $I \propto n(\omega)$  and angular frequency  $\omega$  tuned to energy difference between two discrete energy levels of the atoms, i.e.  $\hbar\omega = E_2 E_1$ .
- Taking into account stimulated absorption, atoms are transferred from level 1 to level 2 at a rate

$$\Gamma_{12} = WN_1 n(\omega)$$

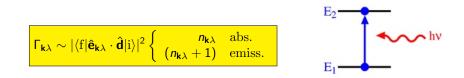
where  $N_1$  atoms in level 1 and W includes matrix elements.

• From spontaneous and stimulated emission processes, the rate of transfer of atoms from level 2 to level 1 is given by

$$\Gamma_{21} = WN_2(n(\omega) + 1)$$

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## Theory of laser: photon equations

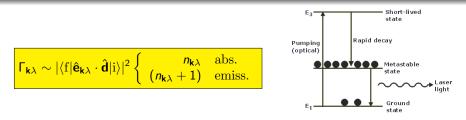


- Since transfer of particles from level 2 to 1 leads to creation of photons in cavity while from 1 to 2 they involve absorption, the rate of change of photon number is given by  $\dot{n} = W(N_2(n+1) N_1n)$ .
- However, to make use of cavity as a photon source, we have to allow photons to leak from the cavity through imperfect mirrors. Taking into account this and other loss processes, we have

$$\dot{n} = DWn + N_2W - rac{n}{ au_{
m ph}}$$

where  $D=N_2-N_1$  denotes population imbalance and  $1/\tau_{\rm ph}$  is the total loss rate.

## Theory of laser: matter equations

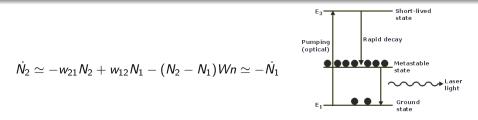


- Without further external processes, photons would escape from cavity and the system would relax into ground state – To create a steady-state photon population, energy must be pumped into the system in the form of excitations.
- Achieved by transferring atoms between 1 and 2 via level 3 by non-resonant optical pump. If lifetime of 3 is short, occupancy is effectively zero, rate of transfer of particles from 2 to 1,

$$N_2 \simeq -w_{21}N_2 + w_{12}N_1 - (N_2 - N_1)Wn \simeq -N_1$$

where we have dropped small contribution from spontaneous emission, and  $w_{12}$ ,  $w_{21}$  denote net non-resonant transition rates.

## Theory of laser: stationary equations



• Without cavity photons (n = 0), since  $N_1 + N_2 \simeq N$ , in steady state,

$$D^{(0)} = N_2^{(0)} - N_1^{(0)} = N \frac{w_{12} - w_{21}}{w_{12} + w_{21}}$$

denotes unsaturated inversion.

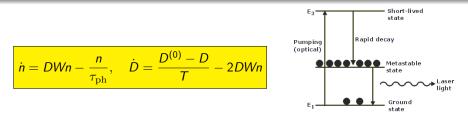
• Restoring the cavity photons, we have

$$\dot{D} = \dot{N}_2 - \dot{N}_1 = \frac{D^{(0)} - D}{T} - 2DWn$$

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where  $1/T = w_{12} + w_{21}$  represents typical relaxation rate.

## Theory of laser: stationary equations



• In steady-state operation,  $\dot{n} = \dot{D} = 0$ , population imbalance

$$D\equiv N_2-N_1=\frac{D^{(0)}}{1+2TWn}$$

From this result, we find the steady state photon number

$$\textit{n} = \frac{\textit{D}^{(0)}\textit{W} - 1/\tau_{\rm ph}}{2\textit{TW}/\tau_{\rm ph}}$$

• This result shows that the system will only start lasing when the unsaturated inversion exceeds a threshold,  $D_{\rm co}^{(0)} > 1/\tau_{\rm ph} W_{\rm constant}$ 

- Although the analysis above addressed the threshold conditions for the laser, it does not provide any insight into the coherence properties of the radiation field.
- In fact, one may show that the radiation field generated by the laser cavity forms a **coherent or Glauber state**.
- The proof of this statement and the coherence properties that follow would take us on a considerable detour see Part III quantum optics.
- However, we can gain some insight into the properties and physical manifestations of coherent states by looking at a toy example; but first we must define what we mean by a coherent state.

• A coherent state is defined as an eigenstate of the annihilation operator,

$$| eta 
angle = eta | eta 
angle$$

#### Since a is not Hermitian, $\beta$ can take complex eigenvalues.

• The eigenstates are constructed from the harmonic oscillator ground state the by action of the unitary operator,

$$|eta
angle=\hat{U}(eta)|0
angle, \qquad \hat{U}(eta)=e^{eta a^{\dagger}-eta^{*}a}, \qquad \hat{U}^{\dagger}(eta)\hat{U}(eta)=\mathbb{I}$$

• The proof follows from the identity (problem set I),

$$a\hat{U}(eta) = \hat{U}(eta)(a+eta), \quad \text{i.e.} \quad a\hat{U}(eta)|0
angle = eta\hat{U}(eta)|0
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• With  $|n\rangle = \frac{1}{\sqrt{n!}} (a^{\dagger})^n |0\rangle$ , we can write  $|\beta\rangle = \sum_n e^{-|\beta|^2/2} \frac{\beta^n}{\sqrt{n!}}$ 

showing that the probability of observing n excitations

$$P_n = |\langle n|\beta\rangle|^2 = e^{-|\beta|^2} \frac{|\beta|^{2n}}{n!}$$

is a Poisson distribution with average occupation,  $\langle \beta | a^{\dagger} a | \beta \rangle = |\beta|^2$ .

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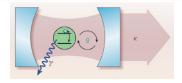
• With  $|n\rangle = \frac{1}{\sqrt{n!}} (a^{\dagger})^n |0\rangle$ , we can write  $|\beta\rangle = \sum_n e^{-|\beta|^2/2} \frac{\beta^n}{\sqrt{n!}} |n\rangle$ 

showing that the probability of observing n excitations

$$P_n = |\langle n|\beta\rangle|^2 = e^{-|\beta|^2} \frac{|\beta|^{2n}}{n!}$$

is a Poisson distribution with average occupation,  $\langle \beta | a^{\dagger} a | \beta \rangle = |\beta|^2$ .

But how can we prepare a system in a coherent state?



• Consider a single two-level atom resonantly coupled to a single cavity mode – the quantum Hamiltonian of the coupled system,

$$\hat{H} = rac{1}{2}\hbar\omega\sigma_z + \hbar\omega\left(a^{\dagger}a + rac{1}{2}
ight) + \hbar g(\sigma_- a + \sigma_+ a^{\dagger})$$

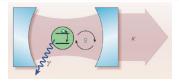
• When excitations of two level system are driven by an external pump, it can behave as a classical dipole source for the cavity mode leading to the driven harmonic oscillator Hamiltonian,

$$\hat{H} \simeq \hat{H}_{\mathrm{rad}} + V(t) = \hbar \omega \left( a^{\dagger} a + \frac{1}{2} \right) + i\hbar \left( f(t) a^{\dagger} - f^{*}(t) a \right)$$

where  $f(t) = f_0 e^{-i\omega t}$ 

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$$\hat{H} = \hbar\omega \left( a^{\dagger}a + \frac{1}{2} \right) + i\hbar \left( f(t)a^{\dagger} - f^{*}(t)a \right)$$
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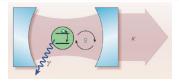
- If photon system is prepared in ground state,  $|0\rangle$ , the perturbation drives the system into a coherent state.
- To understand how, let us turn to the interaction representation,  $i\hbar\partial_t|\psi(t)\rangle_{\rm I} = V_{\rm I}|\psi(t)\rangle_{\rm I}$  where  $|\psi(t)\rangle_{\rm I} = e^{i\hat{H}_0t/\hbar}|\psi(t)\rangle_{\rm S}$ . With  $e^{i\omega ta^{\dagger}a}ae^{-i\omega ta^{\dagger}a} = e^{-i\omega t}a$ ,

$$V_{\rm I}(t) = e^{i\hat{H}_0t/\hbar}i\hbar\left(f(t)a^{\dagger} - f^*(t)a\right)e^{-i\hat{H}_0t/\hbar} = i\hbar\left(f_0a^{\dagger} - f_0^*a\right)$$

• Since  $V_{I}(t)$  is time-independent, the time-evolution operator, defined by the equation  $i\hbar\partial_{t}U_{I}(t) = V_{I}U_{I}(t)$ , is given simply by

$$U_{\rm I}(t) = \exp\left[(f_0 a^{\dagger} - f_0^* a)t\right]$$

$$\hat{H} = \hbar\omega \left( a^{\dagger}a + \frac{1}{2} \right) + i\hbar \left( f(t)a^{\dagger} - f^{*}(t)a \right)$$
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• Therefore, if the system was prepared in the ground state  $|0\rangle$  at t = 0, at later times,

$$|\psi(t)
angle_{
m I}=\exp[(f_{0}a^{\dagger}-f_{0}^{*}a)t]|0
angle=e^{-|f_{0}|^{2}t^{2}/2}e^{f_{0}a^{\dagger}t}|0
angle$$

Reexpressed in the Schrödinger representation,

$$|\psi(t)
angle_{
m S}=e^{-i\hat{H}_0t/\hbar}|\psi(t)
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m I}=e^{-|f_0|^2t^2/2}e^{f_0e^{-i\omega t}a^{\dagger}t}|0
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 A classical oscillatory force drives a system prepared in the vacuum state into a coherent state with an excitation number which climbs as |f<sub>0</sub>|<sup>2</sup>t<sup>2</sup>.

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#### Field theory: from phonons to photons:

From particles to fields: classical field theory of harmonic atomic chain; quantization of atomic chain; phonons. Classical theory of the EM field; waveguide; quantization of the EM field and photons.

#### Time-dependent perturbation theory:

Rabi oscillations in two level systems; perturbation series; sudden approximation; harmonic perturbations and Fermi's Golden rule.

### Radiative transitions:

Light-matter interaction; spontaneous emission; absorption and stimulated emission; Einstein's A and B coefficents; dipole approximation; selection rules; <sup>†</sup>lasers.

### Scattering theory

Elastic scattering; cross section; method of particle waves; Born approximation; scattering of identical particles.

### 6 Relativistic quantum mechanics:

Klein-Gordon equation; Dirac equation; relativistic covariance and spin; free relativistic particles and the Klein paradox; antiparticles; coupling to EM field: minimal coupling and the connection to non-relativistic quantum mechanics; <sup>†</sup>field quantization.