## Theoretical Physics 1 Answers to Examination 2005

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Q1. Bookwork: Hamilton's principle is $\delta \int \mathrm{d} t L\left(q_{i}, \dot{q}_{i}, t\right)=0$ and leads (via the calculus of variations) to

$$
\begin{equation*}
\frac{\mathrm{d}}{\mathrm{~d} t} \frac{\partial L}{\partial \dot{q}_{i}}=\frac{\partial L}{\partial q_{i}} \tag{1}
\end{equation*}
$$

i.e. $N 2$ nd-order equations for the coordinates $q_{i}$.

The Lagrangian is

$$
\begin{equation*}
L=T-V=\frac{m}{2}\left(l^{2} \dot{\theta}^{2}+l^{2} \omega^{2} \sin ^{2} \theta\right)+m g l \cos \theta \tag{2}
\end{equation*}
$$

Evaluating the Euler-Lagrange equation

$$
\begin{equation*}
\frac{d}{d t}\left(\frac{\partial L}{\partial \dot{\theta}}\right)=\frac{\partial L}{\partial \theta} \tag{3}
\end{equation*}
$$

gives

$$
\begin{equation*}
m l^{2} \ddot{\theta}=-m g l \sin \theta+m l^{2} \omega^{2} \sin \theta \cos \theta \tag{4}
\end{equation*}
$$

For small oscillations around $\theta=0$ this may be rewritten as

$$
\begin{equation*}
m l^{2} \ddot{\theta}=-m g l \theta+m l^{2} \omega^{2} \theta \tag{5}
\end{equation*}
$$

For stability this requires that

$$
\begin{equation*}
m l^{2} \omega^{2}>m g l \tag{6}
\end{equation*}
$$

The rotation rate for which $\theta=0$ is no-longer stable is then

$$
\begin{equation*}
\omega_{C}=\sqrt{\frac{g}{l}} \tag{7}
\end{equation*}
$$

For the stable point with $\theta>0$ at frequencies $\omega>\omega_{C}$ we assume that the system performs small oscillations around the angle $\theta_{0}$ so that $\theta=\theta_{0}+\delta$. Substituting this into equation 4 we find

$$
\begin{equation*}
l \ddot{\delta}=\left(-g \sin \theta_{0}+l \omega^{2} \sin \theta_{0} \cos \theta_{0}\right)+\left(-g \cos \theta_{0}+l \omega^{2}\left(\cos ^{2} \theta-\sin ^{2} \theta\right)\right) \delta \tag{8}
\end{equation*}
$$

Simple harmonic motion only occurs if

$$
\begin{equation*}
g \sin \theta_{0}-m l \omega^{2} \sin \theta_{0} \cos \theta_{0}=0 \tag{9}
\end{equation*}
$$

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Since we know that $0<\theta_{0}<\pi$ then $\sin \theta_{0}>0$ and therefore from Eqn. 9

$$
\begin{equation*}
\cos \theta_{0}=\frac{g}{l \omega^{2}} \tag{10}
\end{equation*}
$$

This reduces Eqn. 8 to

$$
\begin{equation*}
l \ddot{\delta}=\left(-g \cos \theta_{0}+l \omega^{2}\left(\cos ^{2} \theta-\sin ^{2} \theta\right)\right) \delta \tag{11}
\end{equation*}
$$

Substituting from Eqn. 10 into Eqn. 11

$$
\begin{equation*}
l \ddot{\delta}=-l \omega^{2}\left(1-\frac{g^{2}}{\omega^{4} l^{2}}\right) \delta \tag{12}
\end{equation*}
$$

Small oscillations around $\theta=\theta_{0}$ therefore have frequency

$$
\begin{equation*}
\Omega=\omega \sqrt{1-\frac{g^{2}}{\omega^{4} l^{2}}} \tag{13}
\end{equation*}
$$

Q2. To write the given Lagrangian in components (with the convention of summation over pairs of repeated indices): $L=\frac{1}{2} a_{i j} \dot{q}_{i} \dot{q}_{j}-V(q)$. Strictly following the definition of the canonical momentum, we obtain

$$
p_{k}=\frac{\partial L}{\partial \dot{q}_{k}}=\frac{1}{2} a_{i j} \delta_{i k} \dot{q}_{j}+\frac{1}{2} a_{i j} \dot{q}_{i} \delta_{j k}=a_{k j} \dot{q}_{j} .
$$

The Hamiltonian is $H=p_{i} \dot{q}_{i}-L$, with all $\dot{q}_{j}$ substituted by $\dot{q}_{j}=a_{j k}^{-1} p_{k}$. This gives, after a little algebra, the required answer

$$
H=\frac{1}{2} a_{i j}^{-1} p_{i} p_{j}+V(q) .
$$

For the particular case of matrix $\boldsymbol{A}$ given in the question you'll have

$$
L=\frac{1}{2}\left(\dot{q}_{1}^{2}+\dot{q}_{3}^{2}+2 q_{1}^{2} \dot{q}_{2}^{2}-2 q_{1} \dot{q}_{2} \dot{q}_{3}\right)+\frac{1}{2} \log q_{1}
$$

This you need, if you prefer not to invert the matrix $a_{i j}$ to write down the Hamiltonian directly. Either way you should obtain

$$
H=\frac{1}{2}\left(p_{1}^{2}+2 p_{3}^{2}+\frac{p_{2}^{2}}{q_{1}^{2}}+\frac{2 p_{2} p_{3}}{q_{1}}\right)-\frac{1}{2} \log q_{1} .
$$

Now write down the Hamilton equations for the components of momentum:

$$
\dot{p}_{1}=-\partial H / \partial q_{1}=\frac{p_{2}^{2}}{q_{1}^{3}}+\frac{p_{2} p_{3}}{q_{1}^{2}}+\frac{1}{2 q_{1}}
$$

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$$
\begin{gathered}
\dot{p}_{2}=-\partial H / \partial q_{2}=0 \\
\dot{p}_{3}=-\partial H / \partial q_{3}=0 .
\end{gathered}
$$

The last two conditions prove that the corresponding components are the constants of motion.

Now we are told that $p_{1}$ is fixed (and equal to zero), so the first of the equations gives the condition

$$
\begin{equation*}
\frac{p_{2}^{2}}{q_{1}^{3}}+\frac{p_{2} p_{3}}{q_{1}^{2}}+\frac{1}{2 q_{1}}=0 . \tag{5}
\end{equation*}
$$

Resolving this to find the required $p_{3}^{2}$, we obtain

$$
p_{3}^{2}=\left(-\frac{2 p_{2}^{2}+q_{1}^{2}}{2 p_{2} q_{1}}\right)^{2}=1+\frac{p_{2}^{2}}{q_{1}^{2}}+\frac{q_{1}^{2}}{4 p_{2}^{2}} .
$$

This has a minimum with respect to either of its variables, $q_{1}$ or $p_{2}$; a sketch would be nice but not necessary.

Q3. First of all, let's write down the Lagrangian in the simplifying case. Now $\left(d x^{0}, d x^{1}\right)=(c d t, d x)$ and

$$
g_{\mu \nu}=\left(\begin{array}{cc}
g(x) & 0 \\
0 & -g(x)
\end{array}\right)
$$

which gives, after multiplication under the root,

$$
L=-m_{0} \sqrt{c^{2} g(x)-\dot{x}^{2} g(x)}=-m_{0} c \sqrt{g} \sqrt{1-v^{2} / c^{2}}
$$

The l.h.s. of the Euler-Lagrange equation will then take the form

$$
\frac{d}{d t}\left(m_{0} \sqrt{g} \frac{\dot{x}}{\sqrt{c^{2}-\dot{x}^{2}}}\right)=\frac{d}{d t}\left(m_{0} v \frac{\sqrt{g}}{\sqrt{c^{2}-v^{2}}}\right)
$$

(the factor following the $m_{0} v$ is therefore denoted as $\Gamma$ in the question. The r.h.s. is

$$
\frac{\partial L}{\partial x}=-m_{0} c \sqrt{1-v^{2} / c^{2}}\left(\frac{1}{2 \sqrt{g}} \frac{\partial g}{\partial x}\right)=-\frac{m_{0}}{\Gamma} \frac{\partial}{\partial x}\left[\frac{1}{2} g(x)\right]
$$

where $\phi$ is the expression in square brackets.
For the general case of $L=-m_{0} \sqrt{g_{\mu \nu} \dot{x}^{\mu} \dot{x}^{\nu}}$ we just need to be careful with components and indices. For the three spatial components of the 4 -vector variable, we'll have in the l.h.s. of the Euler-Lagrange equation:

$$
\frac{d}{d t}\left(\frac{\partial L}{\partial \dot{x}_{i}}\right)=\frac{d}{d t}\left(-m_{0} \frac{2 g_{i \mu} \dot{x}^{\mu}}{2 \sqrt{g_{\mu \nu} \dot{x}^{\mu} \dot{x}^{\nu}}}\right) \equiv \frac{d}{d t}\left(\gamma g_{i \mu} \dot{x}^{\mu}\right)
$$

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Here $i=(1,2,3)$ and $\mu, \nu=(0,1,2,3)$. Now evaluating the derivatives in the r.h.s. we should group terms together into $\gamma=-m_{0} / \sqrt{g_{\mu \nu} \dot{x}^{\mu} \dot{x}^{\nu}}$ (or, equivalently, without $m_{0}$ as this cancels on both sides of the linear equation):

$$
\frac{\partial L}{\partial x_{i}}=-m_{0} \frac{\left(\partial g_{\mu \nu} / \partial x_{i}\right) \dot{x}^{\mu} \dot{x}^{\nu}}{2 \sqrt{g_{\mu \nu} \dot{x}^{\mu} \dot{x}^{\nu}}} \equiv \frac{1}{2} \gamma\left(\frac{\partial g_{\mu \nu}}{\partial x_{i}}\right) \dot{x}^{\mu} \dot{x}^{\nu}
$$

Q4. Cauchy theorem says

$$
\begin{equation*}
\oint_{C} \mathrm{~d} z f(z)=2 \pi i \sum \text { (residues) } \tag{14}
\end{equation*}
$$

with the counterclockwise closed contour $C$. This is proved by expanding $f(z)$ in a Laurent series about a singular point $z_{0}$

$$
\begin{equation*}
f(z)=\sum_{n=-\infty}^{\infty} f_{n}\left(z-z_{0}\right)^{n} \tag{15}
\end{equation*}
$$

and showing that only the $f_{-1}$ term contributes (proof will not be required). The solution of each of the three integrals is based on noticing that the denominator is a quadratic of a quadratic. The first integral has double poles at $z= \pm i$. These are found easily because the denominator is a quadratic in $x^{2}$. We convert to a closed contour by completion in (say) the upper half-plane.

$$
\begin{equation*}
\operatorname{Res}(x=i)=\lim _{x \rightarrow i} \frac{1}{(2-1)!} \frac{\mathrm{d}}{\mathrm{~d} x}(x-i)^{2} \frac{1}{(x-i)^{2}(x+i)^{2}}=-2(2 i)^{-3} \tag{16}
\end{equation*}
$$

This imples that the integral is

$$
\begin{equation*}
I=2 \pi i .-2(2 i)^{-3}=\frac{\pi}{2} \tag{17}
\end{equation*}
$$

The third integral also has a quadratic form for the denominator. It may be be rewritten

$$
\begin{equation*}
(1+x)^{2}+(1+1 / x)^{2}+1=(x+1 / x)^{2}+2(x+1 / x)+1=(x+1 / x+1)^{2}=0 \tag{18}
\end{equation*}
$$

The solutions are therefore

$$
\begin{equation*}
x=-\exp \left( \pm i \frac{\pi}{3}\right) \tag{19}
\end{equation*}
$$

As before each root is doubly degenerate. We close the contour in the u.h.p

$$
\begin{equation*}
\operatorname{Res}(x=i)=\lim _{x \rightarrow-\exp \left(-i \frac{\pi}{3}\right)} \frac{1}{(2-1)!} \frac{\mathrm{d}}{\mathrm{~d} x} \frac{x^{2}}{\left(x+\exp \left(i \frac{\pi}{3}\right)\right)^{2}}=-\frac{2}{9} \sqrt{3} i \tag{20}
\end{equation*}
$$

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The integral then becomes

$$
\begin{equation*}
I=\frac{4}{9} \sqrt{3} \pi \tag{21}
\end{equation*}
$$

The second integral should be rewritten with the substitution $z=\exp (i \theta)$ so that $\mathrm{d} z=i z \mathrm{~d} \theta$. The cosine terms should be rewritten $\cos \theta=(z+1 / z)^{2}$. A contour around the unit circle is used

$$
\begin{equation*}
I=\oint \frac{1}{\frac{1}{4}(z+1 / z)^{2}+\frac{1}{2}(z+1 / z)+1} \cdot \frac{\mathrm{~d} z}{i z} \tag{22}
\end{equation*}
$$

The denominator is a quadratic that has the solution

$$
\begin{equation*}
z+\frac{1}{z}=-1-p \sqrt{3} i \quad p= \pm 1 \tag{23}
\end{equation*}
$$

This is a second quadratic which has the solution

$$
\begin{equation*}
z=\frac{1+p \sqrt{3} i}{2}+\frac{q}{2} \sqrt{(1+p \sqrt{3} i)^{2}-4} \quad q= \pm 1 \tag{24}
\end{equation*}
$$

The term in the square root can be rewritten

$$
\begin{equation*}
\left.(1+p \sqrt{3} i)^{2}-4\right)=-4 \sqrt{3}\left(\frac{\sqrt{3}}{2}-p \frac{i}{2}\right)=-4 \sqrt{3} \exp \left(-i p \frac{\pi}{6}\right) \tag{25}
\end{equation*}
$$

The poles therefore occur at

$$
\begin{equation*}
z=\frac{1+p \sqrt{3} i}{2}+q 3^{1 / 4} i \exp \left(-i p \frac{\pi}{12}\right) \tag{26}
\end{equation*}
$$

These may be evaluated on a calculator to be $z=0.159+0.4052 i, 0.1593-0.4052 i, 0.840+2.173 i, 0.840-2.173 i$.
Evaluation is considerably aided by noticing that if $z$ is a solution then so are $z^{*}, 1 / z, 1 / z^{*}$. Only the first two poles are in the unit circle. The integral is then

$$
\begin{equation*}
I=2 \pi i\left(\operatorname{Res}\left(z_{1}\right)+\operatorname{Res}\left(z_{2}\right)\right)=\pi \frac{3^{1 / 4} \sqrt{2}}{3}(1+\sqrt{3}) \tag{27}
\end{equation*}
$$

Q5. In the opening "essay part" must mention that a potential flow (that is, a potential $\phi(\boldsymbol{x}, t)$ exists such that $\boldsymbol{v}=\nabla \phi)$ requires that $\operatorname{curl} \boldsymbol{v}=0$. This means zero vorticity - as a consequence: (a) no viscous dissipation, (b) Bernoulli equation applicable. Incompressible fluid then satisfies $\nabla^{2} \phi=0$.

Looking at the circular sandbank from above, there are two separate regions, in which we must solve the Laplacian condition $\nabla^{2} \phi=0$, for $\phi_{1}$ inside (over the bank) and for $\phi_{2}$ outside the bank. 2D polar coordinates are recommended in the question, which means ignoring the vorticity region around the edge of the bank.

The boundary conditions are (require 2 for each of $\phi_{1,2}$ ): At $r \rightarrow \infty$ the radial component $\partial_{r} \phi(2) v_{r}(2)=u_{0} \cos \theta$ for the uniform flow. At $r \rightarrow 0$ we will require no singularity in the solution $\phi_{1}$ (see below). At the interface, $r=a$, we want to match the potentials, $\phi_{1}=\phi_{2}$ (which is a similar level of approximation as ignoring the $z$-nonuniformity near the edge). Finally, we need to consider the mass conservation, i.e. the water flowing into the bank must have the same volume as that flowing over it: matching the flow rate $Q_{1}=v_{r}(1) \cdot[$ area $]=v_{r}(1)[2 \pi a d / 2]$ inside, with the flow rate $Q_{2}=v_{r}(2) \cdot[$ area $]=v_{r}(2)[2 \pi a d]$ outside (where the depth $d$ is twice that over the bank). This gives $\partial_{r} \phi_{1}=2 \partial_{r} \phi_{2}$ at $r=a$.
The solution of the Laplacian in 2D polars should be well known to you, as the multipole expansion:

$$
\phi_{1,2}=a_{0} \ln r+\sum_{n=0}^{\infty}\left(a_{n} r^{n}+\frac{b_{n}}{r^{n}}\right) \cos n \theta .
$$

(This format is equivalent, but easier than using formal Legendre polynomials.)
Now, outside the bank, at $r \rightarrow \infty$, we have

$$
v_{r}(2)=\left.\partial_{r} \phi_{2}\right|_{r \rightarrow \infty}=\sum_{n=0}^{\infty} n a_{n} r^{n-1} \cos n \theta
$$

Matching this with the required $u_{0} \cos \theta$ only leaves the mode $n=1$, with $a_{1}=u_{0}$. Since we'll need to match $\phi_{1}=\phi_{2}$, the requirement of only single harmonic ( $n=1$ ) applies to the inside as well. So,

$$
\phi_{1}=\left(A_{1} r+\frac{B_{1}}{r}\right) \cos \theta ; \quad \phi_{2}=\left(u_{0} r+\frac{b_{1}}{r}\right) \cos \theta
$$

Now, at $r=0$ we don't want the singularity in $\phi_{1}$ and, therefore, $B_{1}=0$. This means $\phi_{1}=A_{1} r \cos \theta$, or the velocity $v_{r}(1)=A_{1} \cos \theta$. This is a uniform flow over the bank! (as in many "dielectric problems" you've seen).
We only need to match the solutions at the bank edge, $r=a$. Here

$$
[\phi:] \quad A_{1} a \cos \theta=\left(u_{0} a+b_{1} / a\right) \cos \theta ; \quad\left[v_{r}:\right] \quad A_{1} \cos \theta=2\left(u_{0}-b_{1} / a^{2}\right) \cos \theta
$$

These are two linear equations for the unknowns $A_{1}, b_{1}$. Resolving them we obtain that $A_{1}$ (which is the value of uniform flow velocity over the bank) is equal to $\frac{4}{3} u_{0}$.

Q6. The Langevin equation can be written for each of the Cartesian components of particle coordinate/velocity:

$$
m \ddot{x}=-\gamma \dot{x}+q E+\xi_{x}(t)
$$

$$
\begin{gather*}
m \ddot{y}=-\gamma \dot{y}+\xi_{y}(t) \\
m \ddot{z}=-\gamma \dot{z}-m g+\xi_{z}(t) \tag{8}
\end{gather*}
$$

with the identical properties of the delta-correlated stochastic force in each direction: $\left\langle\xi_{i}\right\rangle=0,\left\langle\xi_{i}^{2}\right\rangle=\Gamma$.
In the overdamped limit we can neglect the inertial (ballistic) component of motion, that is set the acceleration to zero $\ddot{\boldsymbol{x}}=0$. This also means losing the memory of the initial condition for the particle velocity and only consider the balance of forces in the l.h.s. of the Langevin equations.
Formal stochastic solution in the $y$-direction, which is not affected by any force, is $y(t)=(1 / \gamma) \int \xi(\tau) d \tau$. So $\langle y\rangle=0$ and $\left\langle y^{2}\right\rangle=\left(\Gamma / \gamma^{2}\right) t$, the basic diffusion.
Solution in the $x$-direction is $x(t)=(1 / \gamma) \int \xi(\tau) d \tau+(q E / \gamma) t$. So $\langle x\rangle=(q E / \gamma) t$, the constant drift velocity. The mean square has the cross-term vanishing: $\left\langle x^{2}\right\rangle=\left(\Gamma / \gamma^{2}\right) t+\langle x\rangle^{2}$, which means the deviations from the drift velocity are diffusive, with $D=\Gamma / \gamma^{2}$.
Strictly, the same applies to the $z$-motion, since it also has the constant force applied. However, since there is a restriction (the impenetrable bottom of the vessel at $z=0$ ), the constant-velocity drift is not going to happen indefinitely. Instead the system will approach the steady state. The equilibrium ( $t$-independent) probability requires writing the kinetic equation. There are several methods suggested in lectures, the most comprehensive (Fokker-Planck) is to convert

$$
\dot{z}=-(m g / \gamma)+(1 / \gamma) \xi \quad \text { into } \quad \partial_{t} f(z, t)=\partial_{z}[(m g / \gamma) f]+\left(\Gamma / 2 \gamma^{2}\right) \partial_{z}^{2} f
$$

The steady state is obtained by setting the r.h.s. to zero and integrating the l.h.s. over $z$ :

$$
\frac{m g}{\gamma} f=-\frac{\Gamma}{2 \gamma^{2}} \partial_{z} f \quad \text { or } \quad \frac{d f}{f}=-\frac{2 \gamma m g}{\Gamma} d z
$$

and obtaining the required Boltzmann distribution

$$
\begin{equation*}
f(z)=\text { const } e^{-m g z / k T} \tag{10}
\end{equation*}
$$

(because $\Gamma=2 \gamma k T$ in the present notations).

